## XXV IFWGP – Madrid, 2016

# DRINFEL'D DOUBLES FOR (2+1)-POINCARÉ GRAVITY

<u>IVÁN GUTIÉRREZ-SAGREDO</u>, ÁNGEL BALLESTEROS AND FRANCISCO J. HERRANZ

Departamento de Física, Universidad de Burgos, Spain

E-mail: igsagredo@ubu.es, angelb@ubu.es, fjherranz@ubu.es

#### Abstract

By starting from the complete classification of three-dimensional Lie bialgebra structures given in [1], all possible Drinfel'd double structures for the Poincaré algebra iso(2, 1) are constructed in a physically adapted basis. The corresponding classical r-matrices are given, together with the resulting pairing between basis generators. We explore the relations between the associated quantum Poincaré groups and (2+1) gravity with vanishing cosmological constant, where the former appear naturally as symmetries of the corresponding non-commutative spacetimes.

## 1 Quantum groups and Poisson-Lie groups in (2+1) gravity

Quantum group symmetries play an important role in the quantisation of gravity. They occur:

- In **Hamiltonian quantisation formalisms** such as combinatorial quantisation and in path integral approaches (state sum models, spin foams).
- In phenomenological approaches in three and higher dimensions, in non-commutative geometry models such as κ-Poincaré models and related 'doubly special relativity' theories.

For a given Lie algebra/group, there are many possible quantum deformations. In (2+1)-gravity, the classical counterpart of quantum groups (Poisson-Lie groups) arise naturally:

- Poisson-Lie (PL) structures on the isometry groups of (2+1) spaces with constant curvature play a relevant role as **phase spaces when (2+1) gravity is considered as a Chern-Simons (CS)** gauge theory [2, 3].
- These PL structures are given by certain classical r-matrices that have to be 'compatible' with the CS formalism, in the sense that the symmetric component of r has to be directly related with the Ad-invariant symmetric bilinear form in the CS action.

Therefore, it seems natural that the appropriate quantum groups in (2+1) gravity should be the quantizations of these admissible *r*-matrices and PL symmetries.

## 2 The (2+1) gravity $\leftrightarrow$ Drinfel'd double relation

But such admissible r-matrices have to be identified. As it has been shown in [4]:

All the classical r-matrices coming from a Drinfel'd double structure of the (2+1)Lorentzian isometry groups -(A)dS and Poincaré- fulfill the Fock-Rosly condition and are

#### 4 Classical *r*-matrices from Drinfel'd doubles

Definition [5]: A 2*d*-dimensional Lie algebra  $\mathfrak{a}$  has the structure of a (classical) Drinfel'd double if there exists a basis  $\{X_1, \ldots, X_d, x^1, \ldots, x^d\}$  of  $\mathfrak{a}$  in which the Lie bracket takes the form

$$[X_i, X_j] = c_{ij}^k X_k \qquad [x^i, x^j] = f_k^{ij} x^k \qquad [x^i, X_j] = c_{jk}^i x^k - f_j^{ik} X_k.$$

As a consequence:

• An Ad-invariant symmetric bilinear form on  $\mathfrak a$  is given by

$$\langle X_i, X_j \rangle = 0$$
  $\langle x^i, x^j \rangle = 0$   $\langle x^i, X_j \rangle = \delta^i_j$   $\forall i, j.$ 

• A quadratic Casimir operator for a is always given by

$$C = \frac{1}{2} \sum_{i} (x^{i} X_{i} + X_{i} x^{i}).$$

• If a is a DD Lie algebra, its corresponding Lie group can be always endowed with a PL structure generated by the canonical classical *r*-matrix

$$r = \sum_{i} x^{i} \otimes X_{i} = \frac{1}{2} \sum_{i} x^{i} \wedge X_{i} + \frac{1}{2} \sum_{i} (x^{i} \otimes X_{i} + X_{i} \otimes x^{i})$$

which is a (constant) solution of the Classical Yang-Baxter equation [[r, r]] = 0.

Since the symmetric component of r coincides with the tensorized form of the canonical quadratic Casimir element in  $\mathfrak{a}$ , then the Fock-Rosly condition is automatically fulfilled.

Therefore, in (2+1)-Poincaré gravity, any DD structure on  $so(2,1) \ltimes \mathbb{R}^3$  will provide an admissible *r*-matrix.

## 5 Admissible Poincaré *r*-matrices from $so(2,1) \ltimes \mathbb{R}^3$ DDs

The complete classification of the six-dimensional DD Lie algebras is known [1]. In particular, the Poincaré Lie algebra  $so(2,1) \ltimes \mathbb{R}^3$  admits only five DD structures  $(\mathfrak{g}, \mathfrak{g}^*)$ :

- 1.  $(so(2,1), \mathfrak{r}_3(1)) \equiv \text{Case 3 from } [1]$
- 2.  $(\mathfrak{r}_3(1), \mathfrak{n}_3) \equiv \text{Case 10 from } [1]$
- 3.  $(\mathfrak{r}'_3(1), \mathfrak{n}_3) \equiv \text{Case 13 from } [1]$
- 4.  $(\mathfrak{r}'_3(1), \mathfrak{s}_3(0)) \equiv \text{Case } 14' \text{ from } [1]$
- 5.  $(\mathfrak{r}_3(-1), \mathfrak{r}'_3(1)) \equiv \text{Case 14 from } [1]$

The corresponding **admissible** *r***-matrices** coming from them are given below. In all cases we give the most general admissible r-matrix.

- 1.  $r'_{A} = \alpha_1 (J \wedge K_1 + K_2 \wedge K_1) + \beta_1 (P_0 \wedge J + K_1 \wedge P_2 + P_1 \wedge K_2)$
- 2.  $r'_{\rm A} = \alpha_2 \left( P_0 \wedge K_2 + P_2 \wedge K_2 \right) + \beta_2 \left( J \wedge P_2 + \frac{1}{2} \left( P_1 \wedge K_2 + J \wedge P_0 + K_1 \wedge P_2 \right) \right)$
- **3.**  $r'_{A} = \alpha_3 \left( P_0 \wedge K_2 + P_2 \wedge K_2 \right) \beta_3 \left( J \wedge P_2 + \frac{1}{2} \left( P_1 \wedge K_2 + J \wedge P_0 + K_1 \wedge P_2 \right) \right) + \delta_3 P_1 \wedge P_0 + \gamma_3 2 \left( P_2 \wedge P_0 \right) + \epsilon_3 2 P_1 \wedge P_2$
- 4.  $r'_{\rm A} = \alpha_4 \left( 2P_1 \wedge J + J \wedge P_0 + K_1 \wedge P_2 + P_1 \wedge K_2 \right)$
- 5.  $r'_{\rm A} = \alpha_5 \left( P_1 \wedge J + \frac{1}{2} \left( J \wedge P_0 + K_1 \wedge P_2 + P_1 \wedge K_2 \right) \right) + \beta_5 P_0 \wedge P_1$

It is worth noting that the second and fourth r-matrices are just particular cases of the third and fifth ones, respectively. This fact will be translated to the resulting noncommutative space-times below.

compatible with the CS formalism.

In this contribution we present in detail the results for the Poincaré case:

- We give the 5 possible Drinfel'd double structures for the Poincaré Lie algebra iso(2,1).
- We obtain 5 candidates for quantum deformations of the Poincaré symmetries that would be appropriate in a (2+1)-gravity setting.
- The new **Poincaré non-commutative spacetimes** coming from such quantum groups should be relevant in the corresponding (2+1) quantum gravity theory.

### 3 The Poincaré isometry group and its Lie algebra

Any solution of the 3d vacuum Einstein equations is of constant curvature which cosmological constant  $\Lambda$  and is locally isometric to one of the six standard spacetimes:

$\Lambda > 0$	$\Lambda = 0$	$\Lambda < 0$
$dS^{2+1} = SO(3,1)/SO(2,1)$ Isom $(dS^{2+1}) = SO(3,1)$	$\begin{split} \mathbf{M}^{2+1} &= ISO(2,1)/SO(2,1)\\ \mathrm{Isom}(\mathbf{M}^{2+1}) &= ISO(2,1) \end{split}$	$AdS^{2+1} = SO(2,2)/SO(2,1)$ Isom $(AdS^{2+1}) = SO(2,2)$
$\mathbf{S}^3 = SO(4)/SO(3)$ Isom $(\mathbf{S}^3) = SO(4)$	$\mathbf{E}^3 = ISO(3)/SO(3)$ Isom( $\mathbf{E}^3$ ) = $ISO(3)$	$\mathbf{H}^3 = SO(3,1)/SO(3)$ Isom $(\mathbf{H}^3) = SO(3,1)$

- Euclidean signature: the three-sphere  $S^3$  ( $\Lambda > 0$ ), 3d hyperbolic space  $H^3$  ( $\Lambda < 0$ ) and 3d Euclidean space  $E^3$  ( $\Lambda = 0$ ).
- Lorentzian signature: the 3d de Sitter space  $dS^{2+1}$  ( $\Lambda > 0$ ) [4], Anti-de Sitter space  $AdS^{2+1}$  ( $\Lambda < 0$ ) [4] and Minkowski space  $M^{2+1}$  ( $\Lambda = 0$ ) [6].

The Lie algebra of the Poincaré isometry group of the (2+1) Minkowski spacetime can be written in a kinematical basis in terms of generators J (rotation),  $K_1, K_2$  (boosts) and  $P_0, P_1, P_2$  (translations), where the latter commute.

The explicit commutation relations are:

$$\begin{bmatrix} J, K_1 \end{bmatrix} = K_2, \qquad \begin{bmatrix} J, K_2 \end{bmatrix} = -K_1, \qquad \begin{bmatrix} K_1, K_2 \end{bmatrix} = -J, \\ \begin{bmatrix} J, P_0 \end{bmatrix} = 0, \qquad \begin{bmatrix} J, P_1 \end{bmatrix} = P_2, \qquad \begin{bmatrix} J, P_2 \end{bmatrix} = -P_1, \\ \begin{bmatrix} K_1, P_0 \end{bmatrix} = P_1, \qquad \begin{bmatrix} K_1, P_1 \end{bmatrix} = P_0, \qquad \begin{bmatrix} K_1, P_2 \end{bmatrix} = 0, \\ \begin{bmatrix} K_2, P_0 \end{bmatrix} = P_2, \qquad \begin{bmatrix} K_2, P_1 \end{bmatrix} = 0, \qquad \begin{bmatrix} K_2, P_2 \end{bmatrix} = P_0, \\ \begin{bmatrix} P_0, P_1 \end{bmatrix} = 0, \qquad \begin{bmatrix} P_0, P_2 \end{bmatrix} = 0, \qquad \begin{bmatrix} P_1, P_2 \end{bmatrix} = 0.$$

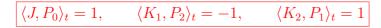
$$(3.1)$$

We have **two quadratic Casimir elements** given by

$$C_1 = P_0^2 - P_1^2 - P_2^2.$$
  

$$C_2 = \frac{1}{2} \left( J P_0 + P_0 J + K_2 P_1 + P_1 K_2 - (K_1 P_2 + P_2 K_1) \right).$$

and associated to the second one, the following non-degenerate symmetric Ad-invariant bilinear form which is the relevant one in (2+1)-gravity:



## 6 Five noncommutative (2+1) Poincaré spacetimes

These five admissible *r*-matrices give rise to five new quantum Poincaré groups, whose explicit construction is in progress [6]. We present here the expressions for the PL structure corresponding to these five cases.

The Poisson brackets that define the PL structure on  $\mathcal{C}^{\infty}(ISO(2,1))$  associated to a given classical *r*-matrix  $r = r^{ij}X_i \otimes X_j$  are given by the **Sklyanin bracket** 

$$\{f,g\} = r^{ij}(X_i^L f X_j^L g - X_i^R f X_j^R g) \qquad f,g \in \mathcal{C}^{\infty}(ISO(2,1))$$

Under a suitable parametrization of the ISO(2,1) group in terms of geodesic parallel coordinates, and after computing the left and right invariant vector fields on it, the Poisson-Lie brackets among the spacetime coordinates read [6]

Case	$\{x_0, x_1\}$	$\{x_0, x_2\}$	$\{x_1, x_2\}$
1.	$-\alpha_1(x_0x_2 + x_1x_2) + 2\beta_1x_2$	$\alpha_1(x_0x_1 + x_1^2) - 2\beta_1x_1$	$\alpha_1(x_0x_1 + x_0^2) - 2\beta_1x_0$
2.	0	$\alpha_2 \left( -x_0 + x_2 \right)$	$\beta_2 \left( -x_0 + x_2 \right)$
3.	$0 + f_1(\xi_1, \xi_2, \theta)$	$\alpha_3 \left( -x_0 + x_2 \right) + f_2(\xi_1, \xi_2, \theta)$	$\beta_3 \left( -x_0 + x_2 \right) + f_3(\xi_1, \xi_2, \theta)$
4.	0	0	$-2\alpha_4(x_0+x_1)$
5.	$0+f_4(\xi_1,\xi_2,\theta)$	$0+f_5(\xi_1,\xi_2,\theta)$	$-2\alpha_5(x_0+x_1)+f_6(\xi_1,\xi_2,\theta)$

Note that:

- The quantization of these algebras will give **new non-commutative Poincaré spacetimes**.
- As stated above, 2 and 4 are particular cases of 3 and 5, respectively. The commutation relations for each case are recovered when the appropriate parameters vanish.
- Cases 1, 2 and 4 present commutation relations among spacetime coordinates where only these same coordinates appear. This fact is in agreement with the idea of a (physical) noncommutative spacetime.

#### Acknowledgments

This work was partially supported by the Spanish MINECO under grant MTM2013-43820-P and by Junta de Castilla y León under grants BU278U14 and VA057U16. I.G.S. acknowledges a predoctoral grant from the European Social Fund and Junta de Castilla y León.

#### References

- [1] X. Gomez, Classification of three-dimensional Lie bialgebras, J. Math. Phys. 41 (2000) 4939.
- [2] A.Y. Alekseev and A.Z. Malkin A Z, Commun. Math. Phys. 169, 99 (1995)
   V.V. Fock and A.A. Rosly, Am. Math. Soc. Transl., 191 67 (1999)
   C. Meusburger and B.J. Schroers, Nucl. Phys. B 806, 462 (2009)
- [3] A. Achucarro and P.K. Townsend, *Phys. Lett. B* 180, 89 (1986)
   E. Witten, *Nucl. Phys. B* 311, 46 (1988)
- [4] A. Ballesteros, F.J. Herranz and C. Meusburger, Class. Quantum Grav. 30, 155012 (2013)
- [5] V.G. Drinfel'd, Proc. Int. Congress of Math., AMS, 798 (1987)
- [6] A. Ballesteros, I. Gutiérrez-Sagredo, F.J. Herranz, Drinfel'd doubles for (2+1) Poincaré gravity, in preparation (2016).