Mass, Scalar Curvature, &

Kähler Geometry, I

Claude LeBrun Stony Brook University

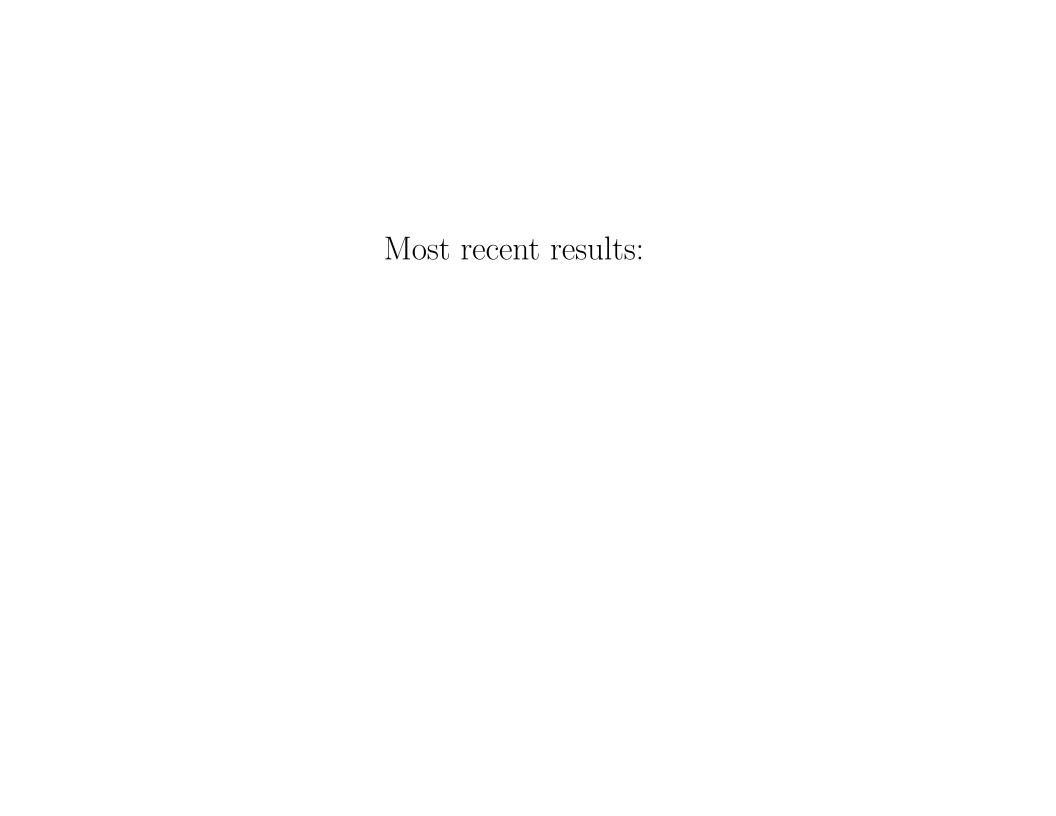
Seminario de Geometría ICMAT, October 29, 2018 Key results joint with

Key results joint with

Hans-Joachim Hein Fordham University Key results joint with

Hans-Joachim Hein Fordham University

Mass in Kähler Geometry Comm. Math. Phys. 347 (2016) 621–653.



Most recent results:

Mass, Kähler Manifolds, and Symplectic Geometry

Most recent results:

Mass, Kähler Manifolds, and Symplectic Geometry

arXiv: 1810.11417 [math.DG]

Recall:

Recall: the scalar curvature of (M^n, g)

$$s:M\to\mathbb{R}$$

is a standard Riemannian invariant that compares the volume of small balls to the Euclidean answer: Recall: the scalar curvature of (M^n, g)

$$s:M\to\mathbb{R}$$

is a standard Riemannian invariant that compares the volume of small balls to the Euclidean answer:

$$\frac{\operatorname{vol}_g(B_{\varepsilon}(p))}{\text{Euclidean answer}} = 1 - s(p) \frac{\varepsilon^2}{6(n+2)} + O(\varepsilon^4)$$



Recall: the scalar curvature of (M^n, g)

$$s:M\to\mathbb{R}$$

is a standard Riemannian invariant that compares the volume of small balls to the Euclidean answer:

$$\frac{\operatorname{vol}_g(B_{\varepsilon}(p))}{\text{Euclidean answer}} = 1 - s(p) \frac{\varepsilon^2}{6(n+2)} + O(\varepsilon^4)$$



The metric g is called scalar-flat if it satisfies $s \equiv 0$.

Similarly, the *Ricci curvature*

 $r: UTM \to \mathbb{R}$

is a standard Riemannian invariant comparing the volume of narrow cones to the Euclidean answer:

Similarly, the *Ricci curvature*

$$r: UTM \to \mathbb{R}$$

is a standard Riemannian invariant comparing the volume of narrow cones to the Euclidean answer:

$$\frac{\operatorname{vol}_g(C_{\varepsilon}(p, v, \Omega))}{\operatorname{Euclidean answer}} \approx 1 - r(v) \frac{n\varepsilon^2}{6(n+2)} + O(\varepsilon^3)$$



Similarly, the *Ricci curvature*

$$r: UTM \to \mathbb{R}$$

is a standard Riemannian invariant comparing the volume of narrow cones to the Euclidean answer:

$$\frac{\operatorname{vol}_g(C_{\varepsilon}(p, v, \Omega))}{\operatorname{Euclidean answer}} \approx 1 - r(v) \frac{n\varepsilon^2}{6(n+2)} + O(\varepsilon^3)$$



The metric g is called Ricci-flat if it satisfies $r \equiv 0$.

On the other hand, the sectional curvature

$$K: Gr_2TM \to \mathbb{R}$$

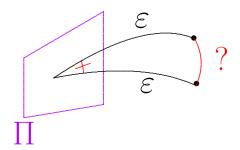
compares distances to the Euclidean answer:

On the other hand, the sectional curvature

$$K: Gr_2TM \to \mathbb{R}$$

compares distances to the Euclidean answer:

$$\frac{\text{length}_g(\text{base of isosceles triangle})}{\text{Euclidean answer}} \approx 1 - K(\Pi) \frac{\varepsilon^2}{6} + O(\varepsilon^3)$$

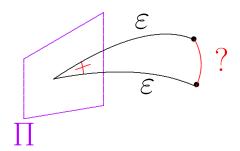


On the other hand, the sectional curvature

$$K: Gr_2TM \to \mathbb{R}$$

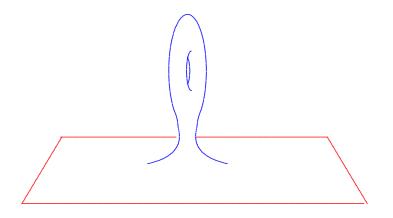
compares distances to the Euclidean answer:

$$\frac{\text{length}_g(\text{base of isosceles triangle})}{\text{Euclidean answer}} \approx 1 - K(\Pi) \frac{\varepsilon^2}{6} + O(\varepsilon^3)$$



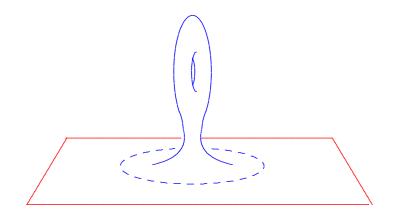
The metric g is called flat if it satisfies $K \equiv 0$.

 (M^n, g) complete non-compact Riemannian n-manifold.



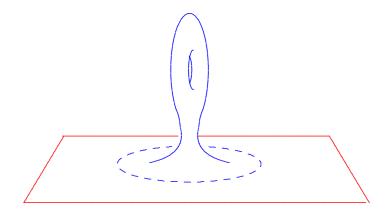
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



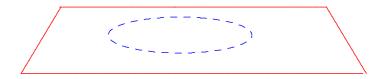
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



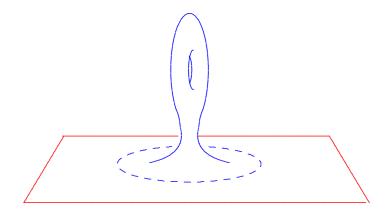
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



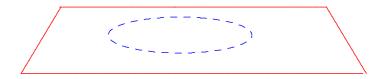
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



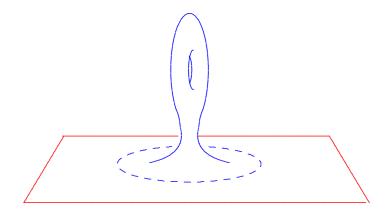
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



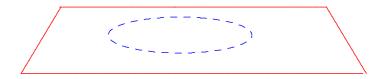
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



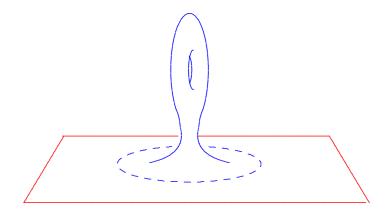
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



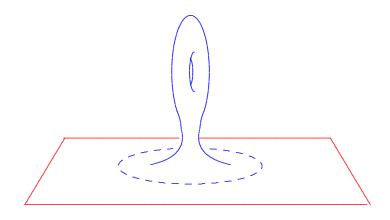
 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

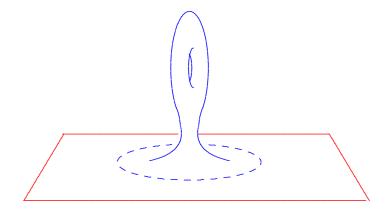


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has curvature ≥ 0 , is it flat?

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

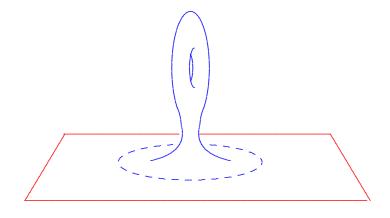


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has sectional curvature ≥ 0 , is it flat?

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

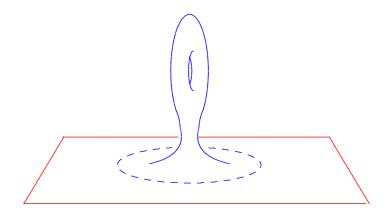


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has sectional curvature ≥ 0 , is it flat? Yes!

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

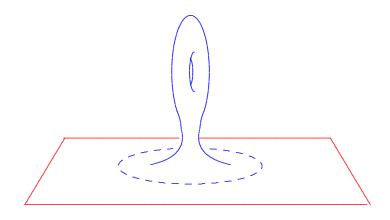


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has sectional curvature ≥ 0 , is it flat? Yes! Distance comparison...

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

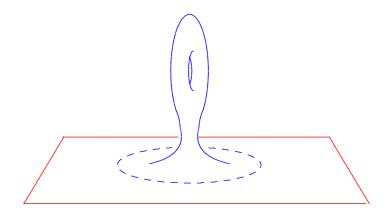


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has curvature ≥ 0 , is it flat?

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

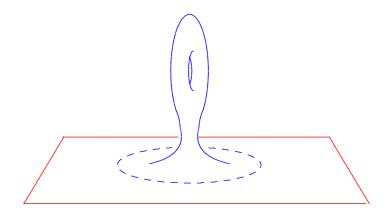


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has Ricci curvature ≥ 0 , is it flat?

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

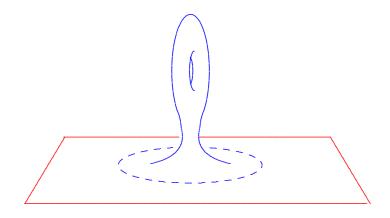


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has Ricci curvature ≥ 0 , is it flat? Yes!

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

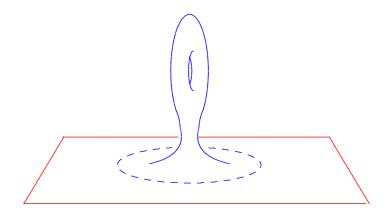


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has Ricci curvature ≥ 0 , is it flat? Yes! Volume comparison...

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

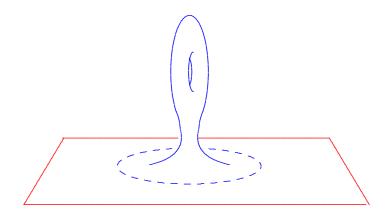


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has curvature ≥ 0 , is it flat?

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

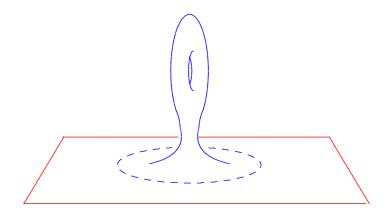


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has scalar curvature ≥ 0 , is it flat?

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact

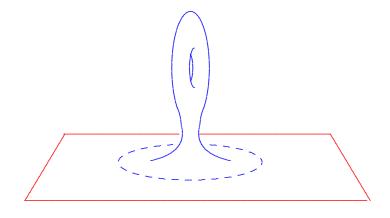


and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has scalar curvature ≥ 0 , is it flat? Yes!

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



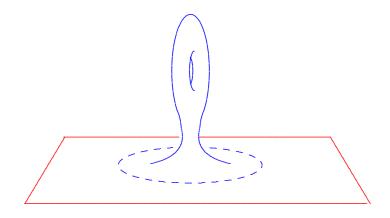
and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has scalar curvature ≥ 0 , is it flat? Yes!

This time, the inspiration comes from physics!

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



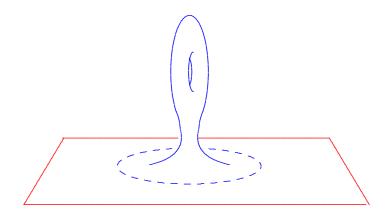
and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has scalar curvature ≥ 0 , is it flat? Yes!

"Positive Mass Theorem"

 (M^n, g) complete non-compact Riemannian n-manifold.

Suppose $\exists K \subset M$ compact



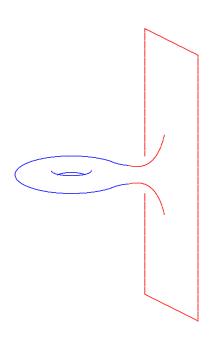
and an isometry $M - K \to \mathbb{R}^n - D^n$. (Euclidean)

If M has scalar curvature ≥ 0 , is it flat? Yes!

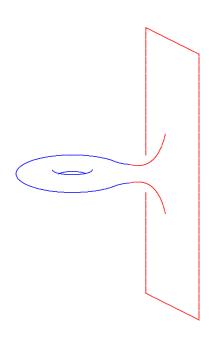
"Positive Mass Theorem"

Get result even with appropriate fall-off to Euclidean. . .

Definition. A complete, non-compact Riemannian n-manifold (M^n, g)

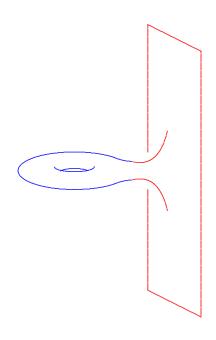


Definition. A complete, non-compact Riemannian n-manifold (M^n, g)



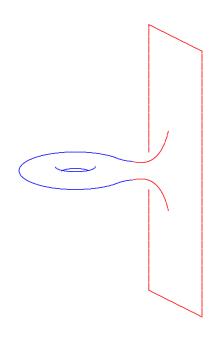
 $n \ge 3$

Definition. A complete, non-compact Riemannian n-manifold (M^n, g) is called asymptotically Euclidean



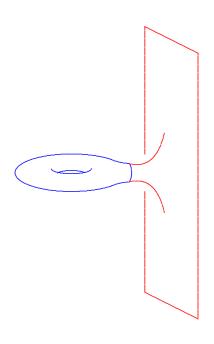
 $n \ge 3$

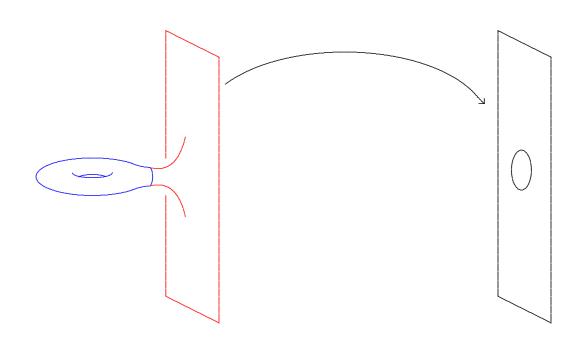
Definition. A complete, non-compact Riemannian n-manifold (M^n, g) is called asymptotically Euclidean (AE)

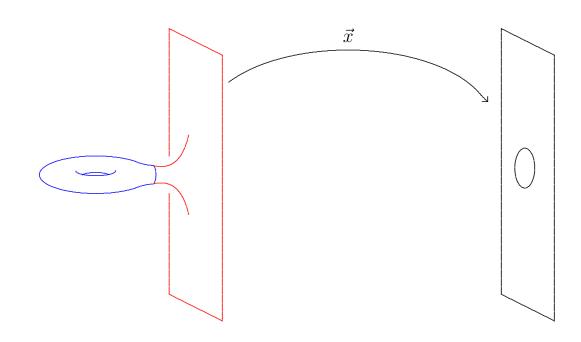


 $n \ge 3$

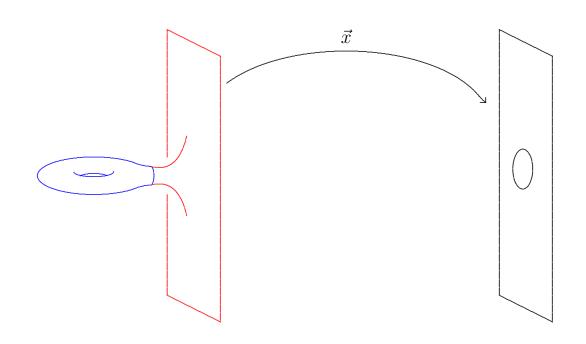
Definition. A complete, non-compact Riemannian n-manifold (M^n, g) is called asymptotically Euclidean (AE) if there is a compact set $K \subset M$



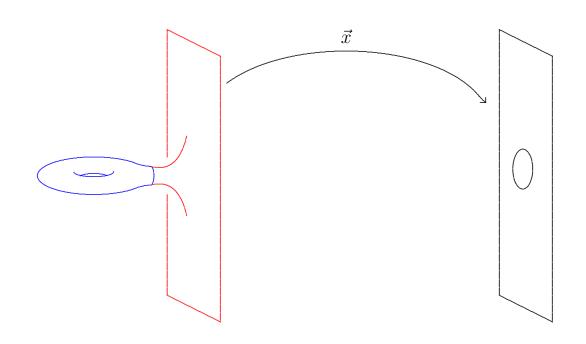




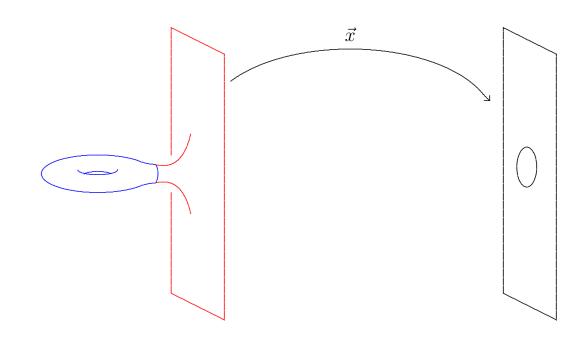
$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$



$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon})$$

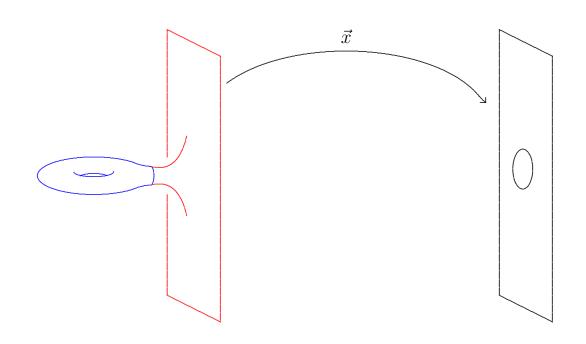


$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad s \in L^1$$

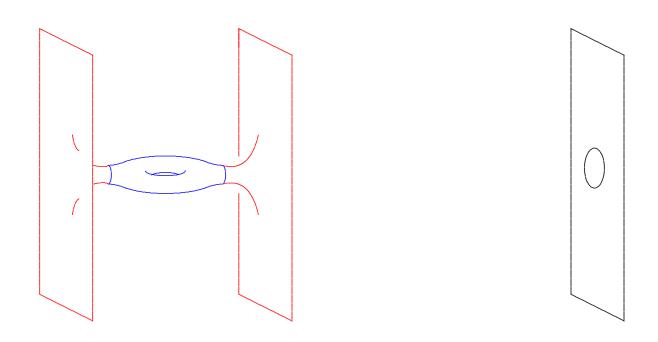


$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$

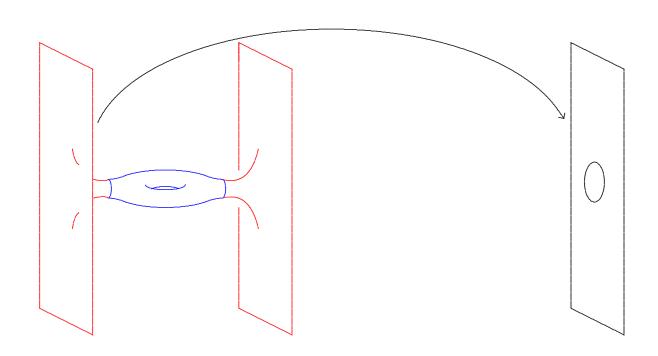
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \text{scalar curvature } \in L^1$$



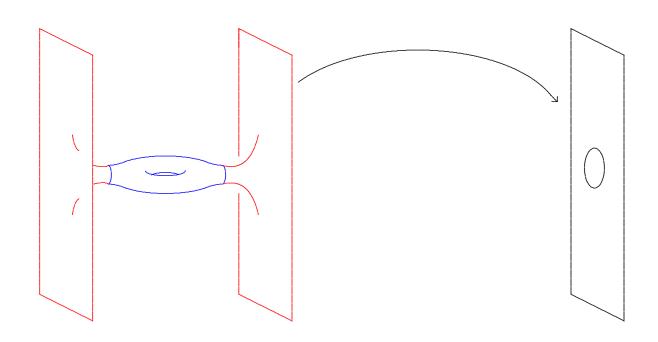
$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad s \in L^1$$



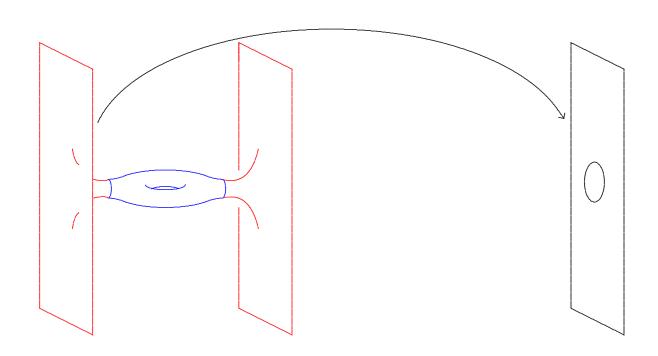
$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$



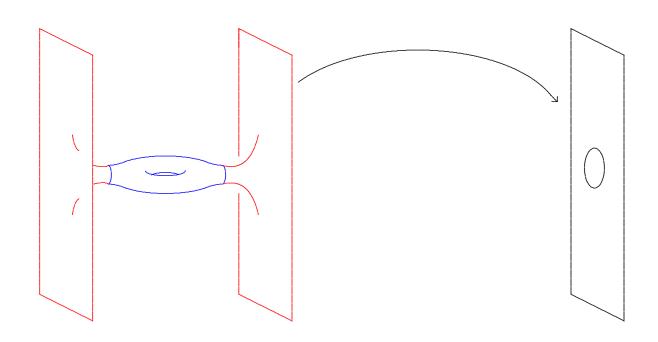
$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$



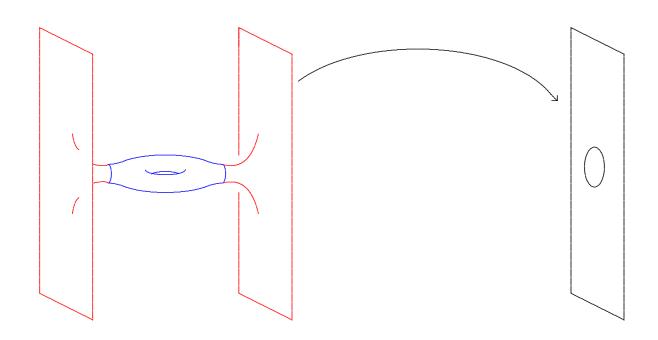
$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$



$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$



$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$



$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$

$$m(M,g) := [g_{ij,i} - g_{ii,j}]$$

$$m(M,g) := [g_{ij,i} - g_{ii,j}] \nu^j$$

$$m(M,g) := \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^{j} \alpha_{E}$$

where

• $\Sigma(\varrho) \approx S^{n-1}$

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

• $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- $\bullet \nu$ is the outpointing Euclidean unit normal;

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

Seems to depend on choice of coordinates!

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

Bartnik/Chruściel (1986):

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

Bartnik/Chruściel (1986): With weak fall-off conditions,

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad s \in L^1$$

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^{j} \alpha_{E}$$

where

- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

Bartnik/Chruściel (1986): With weak fall-off conditions, the mass is well-defined

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

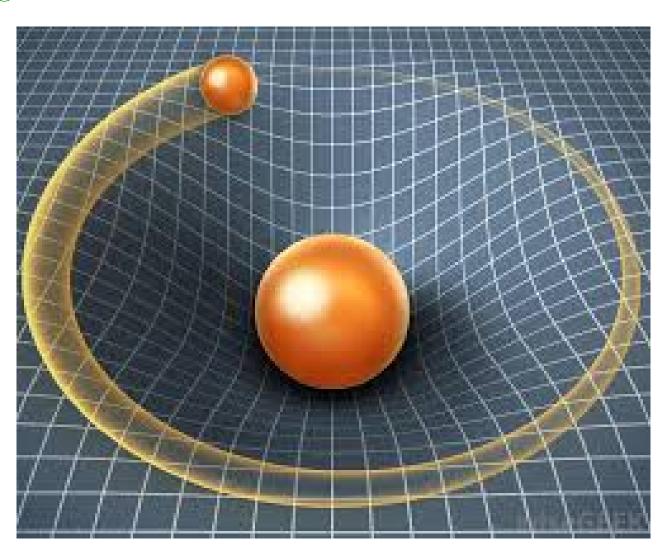
- $\Sigma(\varrho) \approx S^{n-1}$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-form induced by the Euclidean metric.

Bartnik/Chruściel (1986): With weak fall-off conditions, the mass is well-defined & coordinate independent.

When n = 3, ADM mass in general relativity.

When n = 3, ADM mass in general relativity.

When n = 3, ADM mass in general relativity.



When n = 3, ADM mass in general relativity.

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

$$g = -\left(1 - \frac{2m}{\varrho^{n-2}}\right)dt^2 + \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1}d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

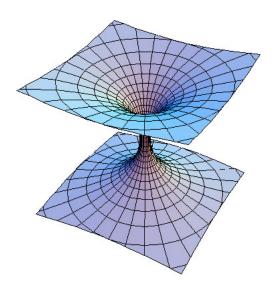
When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Two such regions fit together to form the wormhole metric.



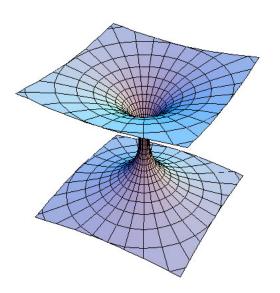
When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Two such regions fit together to form the wormhole metric. Scalar-flat, AE, two ends.



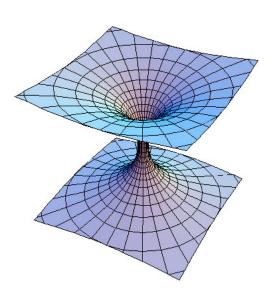
When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Two such regions fit together to form the wormhole metric. Scalar-flat, AE, two ends. Not Ricci-flat, but conformally flat.



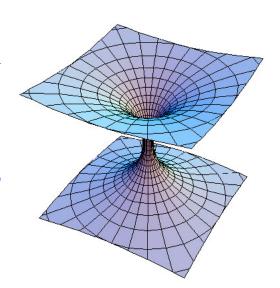
When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Two such regions fit together to form the wormhole metric. Scalar-flat, AE, two ends. Not Ricci-flat, but conformally flat. Same mass m at both ends:



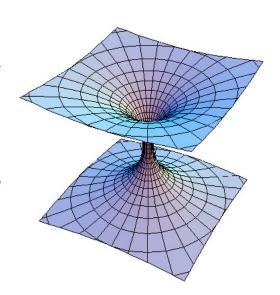
When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Two such regions fit together to form the wormhole metric. Scalar-flat, AE, two ends. Not Ricci-flat, but conformally flat. Same mass m at both ends: "size of throat."



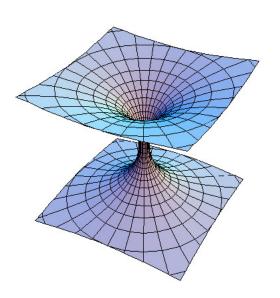
When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 + \frac{m/2}{r^{n-2}}\right)^{4/(n-2)} \left[\sum (dx^j)^2\right]$$

Two such regions fit together to form the wormhole metric. Scalar-flat, AE, two ends. Not Ricci-flat, but conformally flat. Same mass m at both ends: "size of throat."



When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

$$g_{jk} = \left(1 + \frac{2m}{(n-2)r^{n-2}}\right)\delta_{jk} + \cdots$$

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Burns metric on \mathbb{C}^2

When n = 3, ADM mass in general relativity.

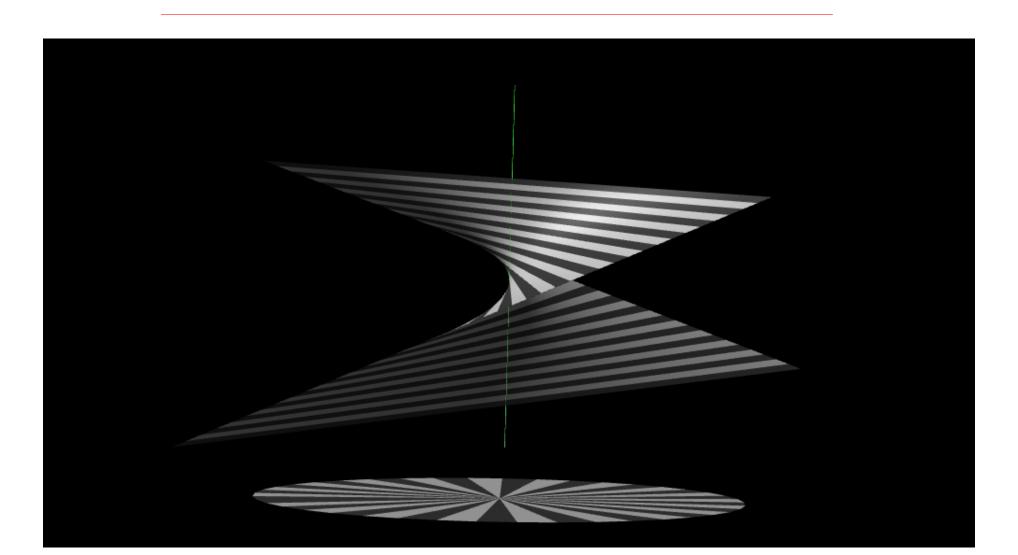
Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Burns metric on $\widetilde{\mathbb{C}^2} \subset \mathbb{C}^2 \times \mathbb{CP}_1$

When n = 3, ADM mass in general relativity.



When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Burns metric on $\widetilde{\mathbb{C}^2} \subset \mathbb{C}^2 \times \mathbb{CP}_1$

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

When n = 3, ADM mass in general relativity.

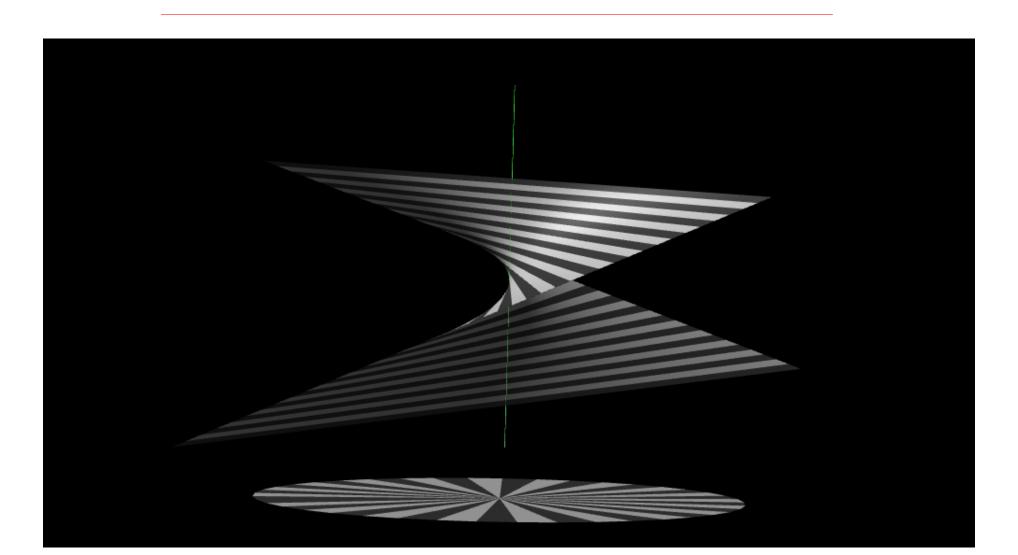
Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Restrict Euclidean × round metric to
$$\begin{vmatrix} z_1 & z_2 \\ \zeta_1 & \zeta_2 \end{vmatrix} = 0.$$

When n = 3, ADM mass in general relativity.



When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Restrict Euclidean × round metric to
$$\begin{vmatrix} z_1 & z_2 \\ \zeta_1 & \zeta_2 \end{vmatrix} = 0$$
.

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

$$\omega = \frac{i}{2}\partial\bar{\partial}\left[u + 3m\log u\right],\,$$

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

$$\omega = \frac{i}{2}\partial\bar{\partial}\left[\mathbf{u} + 3m\log\mathbf{u}\right], \qquad \mathbf{u} = |\mathbf{z}_1|^2 + |\mathbf{z}_2|^2$$

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Scalar-flat-Kähler Burns metric on $\mathbb{C}^2 \subset \mathbb{C}^2 \times \mathbb{CP}_1$:

$$\omega = \frac{i}{2}\partial\bar{\partial}\left[\mathbf{u} + 3m\log\mathbf{u}\right], \qquad \mathbf{u} = |z_1|^2 + |z_2|^2$$

also has mass m.

When n = 3, ADM mass in general relativity.

Reads off "apparent mass" from strength of the gravitational field far from an isolated source.

In any dimension, reproduces "mass" of t=0 hypersurface in (n+1)-dimensional Schwarzschild

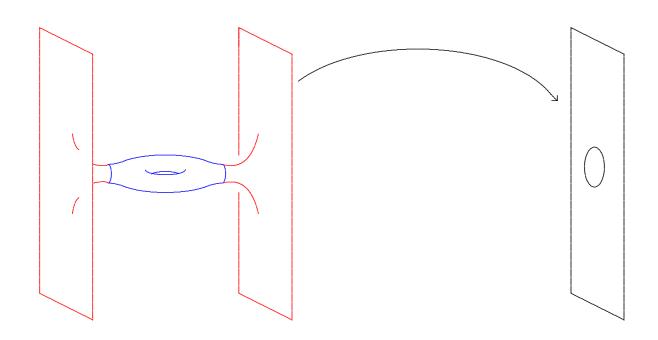
$$g = \left(1 - \frac{2m}{\varrho^{n-2}}\right)^{-1} d\varrho^2 + \varrho^2 h_{S^{n-1}}$$

Scalar-flat-Kähler Burns metric on $\mathbb{C}^2 \subset \mathbb{C}^2 \times \mathbb{CP}_1$:

$$\omega = \frac{i}{2}\partial\bar{\partial}\left[\mathbf{u} + 3\mathbf{m}\log\mathbf{u}\right], \qquad \mathbf{u} = |z_1|^2 + |z_2|^2$$

also has mass m. Again measures "size of throat."

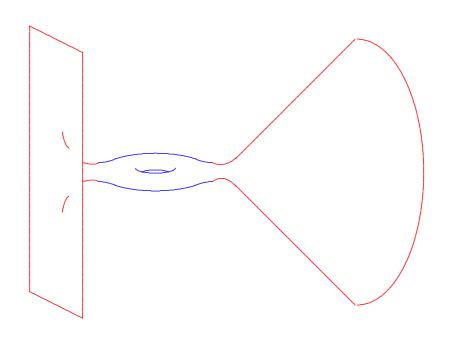
Definition. A complete, non-compact Riemannian n-manifold (M^n, g) is called asymptotically Euclidean (AE) if there is a compact set $K \subset M$ such that each "end" is diffeomorphic to $\mathbb{R}^n - D^n$ in such a manner that



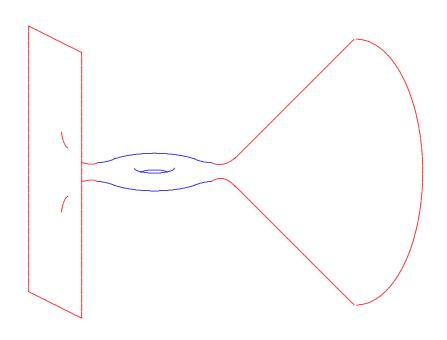
$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$

Interesting generalization...

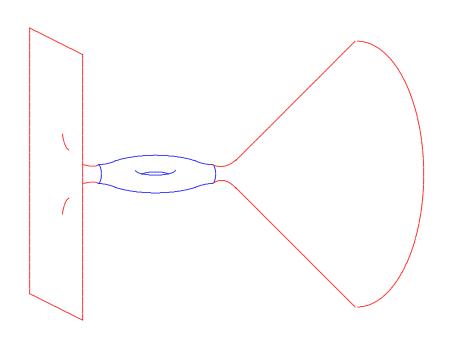
Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean



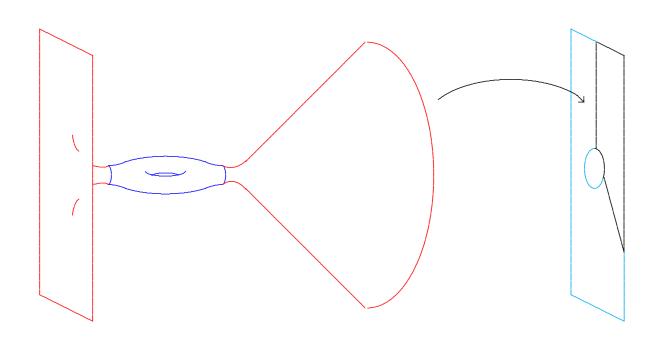
Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE)



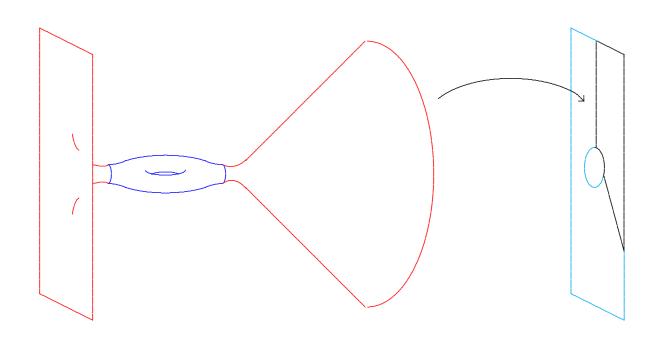
Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$



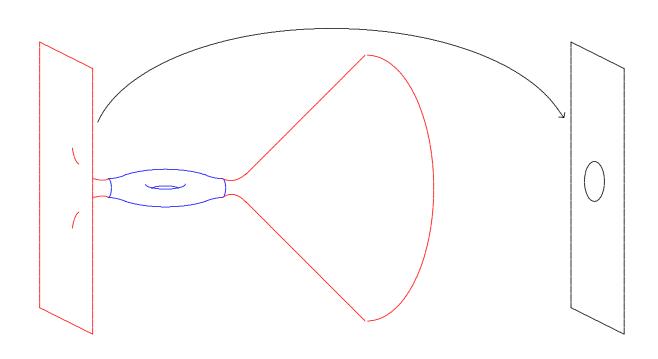
Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$ such that $M - K \approx \coprod_i (\mathbb{R}^n - D^n)/\Gamma_i$,



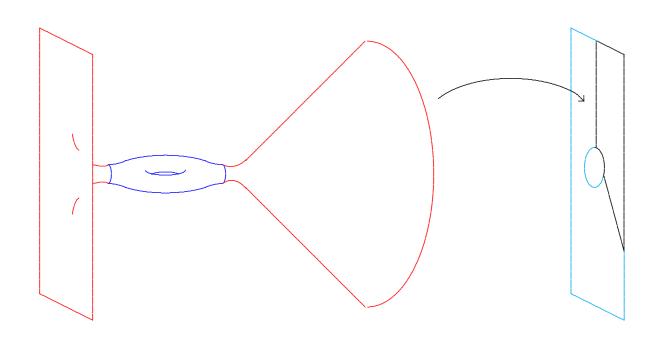
Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$ such that $M - K \approx \coprod_i (\mathbb{R}^n - D^n)/\Gamma_i$, where $\Gamma_i \subset \mathbf{O}(\mathbf{n})$,



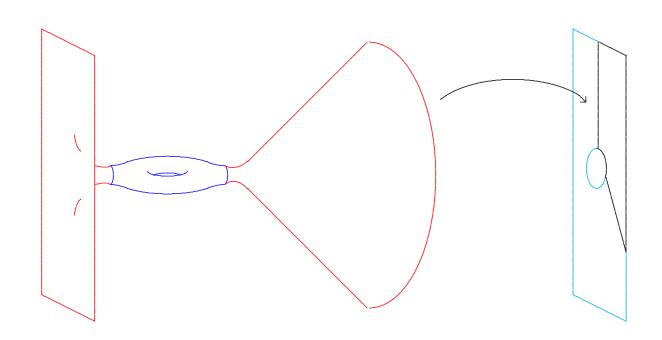
Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$ such that $M - K \approx \coprod_i (\mathbb{R}^n - D^n)/\Gamma_i$, where $\Gamma_i \subset \mathbf{O}(\mathbf{n})$,



Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$ such that $M - K \approx \coprod_i (\mathbb{R}^n - D^n)/\Gamma_i$, where $\Gamma_i \subset \mathbf{O}(\mathbf{n})$,



Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$ such that $M - K \approx \coprod_i (\mathbb{R}^n - D^n)/\Gamma_i$, where $\Gamma_i \subset \mathbf{O}(\mathbf{n})$, such that



$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$

Why consider ALE spaces?

Gravitational Instantons:

Gravitational Instantons:

Bubbling modes for sequences of Einstein metrics.

Gravitational Instantons:

Bubbling modes for sequences of Einstein metrics.

ALE scalar-flat Kähler surfaces:

Gravitational Instantons:

Bubbling modes for sequences of Einstein metrics.

ALE scalar-flat Kähler surfaces:

Bubbling modes for extremal Kähler metrics.

Gravitational Instantons:

Bubbling modes for sequences of Einstein metrics.

ALE scalar-flat Kähler surfaces:

Bubbling modes for extremal Kähler metrics.

Theory used to construct compact Einstein 4-manifolds.

Gravitational Instantons:

Bubbling modes for sequences of Einstein metrics.

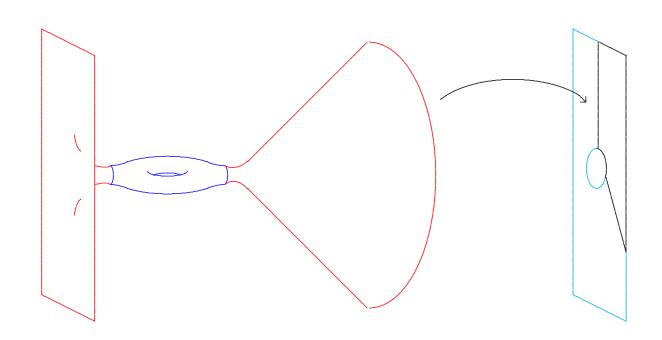
ALE scalar-flat Kähler surfaces:

Bubbling modes for extremal Kähler metrics.

Theory used to construct compact Einstein 4-manifolds.

Will discuss some examples in next lecture.

Definition. Complete, non-compact n-manifold (M^n, g) is asymptotically locally Euclidean (ALE) if \exists compact set $K \subset M$ such that $M - K \approx \coprod_i (\mathbb{R}^n - D^n)/\Gamma_i$, where $\Gamma_i \subset \mathbf{O}(\mathbf{n})$, such that



$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad \mathbf{s} \in L^1$$

Mass still meaningful in this context...

$$m(M,g) := \left[g_{ij,i} - g_{ii,j} \right]$$

$$m(M,g) := [g_{ij,i} - g_{ii,j}] \nu^j$$

$$m(M,g) := \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

$$m(M,g) := \lim_{\varrho \to \infty} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

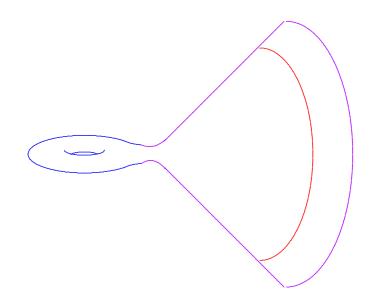
$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

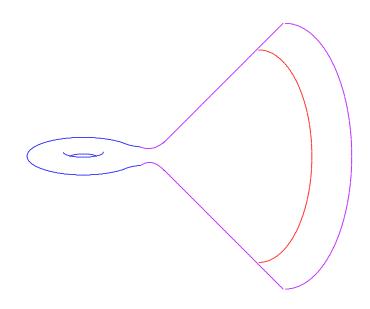
•
$$\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$$



$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

• $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;



$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- $\bullet \nu$ is the outpointing Euclidean unit normal;

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Bartnik/Chruściel (1986):

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Bartnik/Chruściel (1986): With weak fall-off conditions,

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad s \in L^1$$

$$\mathbf{m}(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Bartnik/Chruściel (1986): With weak fall-off conditions, the mass is well-defined

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Bartnik/Chruściel (1986): With weak fall-off conditions, the mass is well-defined & coordinate independent.

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Chruściel-type fall-off:

$$g_{jk} - \delta_{jk} \in C^1_{-\tau}, \quad \tau > \frac{n-2}{2}$$

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Bartnik-type fall-off:

$$g_{jk} - \delta_{jk} \in W^{2,q}_{-\tau}, \quad \tau > \frac{n-2}{2}, \quad q > n$$

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Bartnik-type fall-off \Longrightarrow

$$g_{jk} - \delta_{jk} \in C^{1,\alpha}_{-\tau}, \quad \tau > \frac{n-2}{2}.$$

$$m(M,g) := \lim_{\varrho \to \infty} \frac{\Gamma(\frac{n}{2})}{4(n-1)\pi^{n/2}} \int_{\Sigma(\varrho)} \left[g_{ij,i} - g_{ii,j} \right] \nu^j \alpha_E$$

where

- $\Sigma(\varrho) \approx S^{n-1}/\Gamma_i$ is given by $|\vec{x}| = \varrho$;
- ullet u is the outpointing Euclidean unit normal; and
- α_E is the volume (n-1)-from induced by the Euclidean metric.

Bartnik/Chruściel (1986): With weak fall-off conditions, the mass is well-defined & coordinate independent.

Positive Mass Conjecture:

Any AE manifold

Any AE manifold with $s \ge 0$

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Physical intuition:

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Physical intuition:

Local matter density ≥ 0

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Physical intuition:

Local matter density $\geq 0 \Longrightarrow \text{total mass} \geq 0$.

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Schoen-Yau 2018:

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Schoen-Yau 2018:

General case in arbitrary dimension n.

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Hawking-Pope 1978:

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Hawking-Pope 1978:

Conjectured true in ALE case, too.

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Hawking-Pope 1978:

Conjectured true in ALE case, too.

L 1986:

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Hawking-Pope 1978:

Conjectured true in ALE case, too.

L 1986:

ALE counter-examples.

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Hawking-Pope 1978:

Conjectured true in ALE case, too.

L 1986:

ALE counter-examples.

Scalar-flat Kähler metrics

Any AE manifold with $s \ge 0$ has $m \ge 0$.

Schoen-Yau 1979:

Proved in dimension $n \leq 7$.

Witten 1981:

Proved for spin manifolds (implicitly, for any n).

Hawking-Pope 1978:

Conjectured true in ALE case, too.

L 1986:

ALE counter-examples.

Scalar-flat Kähler metrics

on line bundles $L \to \mathbb{CP}_1$ of Chern-class ≤ -3 .

Scalar-flat Kähler case?

Scalar-flat Kähler case?

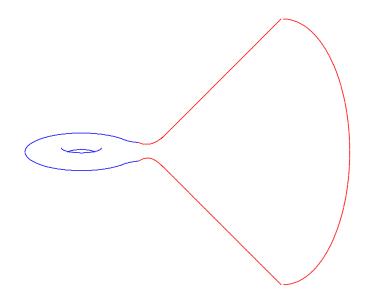
Lemma.

Scalar-flat Kähler case?

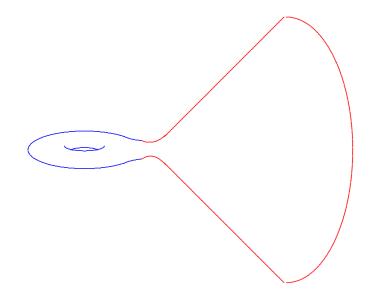
Lemma. Any ALE Kähler manifold

Scalar-flat Kähler case?

Scalar-flat Kähler case?



Scalar-flat Kähler case?



$$n = 2m \ge 4$$

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Several different proofs are known.

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Several different proofs are known.

Several are analytic:

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Several different proofs are known.

Several are analytic:

each end is pseudo-convex at infinity.

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Several different proofs are known.

Several are analytic:

each end is pseudo-convex at infinity.

Another is more topological:

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Several different proofs are known.

Several are analytic:

each end is pseudo-convex at infinity.

Another is more topological:

intersection form on H^2 of compactification.

Scalar-flat Kähler case?

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Upshot:

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Upshot:

Mass of an ALE Kähler manifold is unambiguous.

Mass of ALE Kähler manifolds?

Scalar-flat Kähler case?

Lemma. Any ALE Kähler manifold has only one end.

Upshot:

Mass of an ALE Kähler manifold is unambiguous.

Does not depend on the choice of an end!

Theorem A.

Theorem A. The mass

Theorem A. The mass of an ALE

Theorem A. The mass of an ALE scalar-flat Kähler manifold

Theorem A. The mass of an ALE scalar-flat Kähler manifold is a topological invariant.

Theorem A. The mass of an ALE scalar-flat Kähler manifold is a topological invariant.

That is, m(M, g, J)

Theorem A. The mass of an ALE scalar-flat Kähler manifold is a topological invariant.

That is, m(M, g, J) is completely determined by

Theorem A. The mass of an ALE scalar-flat Kähler manifold is a topological invariant.

That is, m(M, g, J) is completely determined by

ullet the smooth manifold M,

Theorem A. The mass of an ALE scalar-flat Kähler manifold is a topological invariant.

That is, m(M, g, J) is completely determined by

- the smooth manifold M,
- the first Chern class $c_1 = c_1(M, J) \in H^2(M)$ of the complex structure, and

Theorem A. The mass of an ALE scalar-flat Kähler manifold is a topological invariant.

That is, m(M, g, J) is completely determined by

- the smooth manifold M,
- the first Chern class $c_1 = c_1(M, J) \in H^2(M)$ of the complex structure, and
- the Kähler class $[\omega] \in H^2(M)$ of the metric.

Theorem A. The mass of an ALE scalar-flat Kähler manifold is a topological invariant.

That is, m(M, g, J) is completely determined by

- ullet the smooth manifold M,
- the first Chern class $c_1 = c_1(M, J) \in H^2(M)$ of the complex structure, and
- the Kähler class $[\omega] \in H^2(M)$ of the metric.

In fact, we will see that there is an explicit formula for the mass in terms of these data! The explicit formula reproduces the mass in cases where it previously had been laboriously computed from the definition.

Corollary, suggested by Cristiano Spotti:

Corollary, suggested by Cristiano Spotti:

Theorem B.

Corollary, suggested by Cristiano Spotti:

Theorem B. Let (M^4, g, J) be an ALE scalar-flat Kähler surface,

Corollary, suggested by Cristiano Spotti:

Theorem B. Let (M^4, g, J) be an ALE scalarflat Kähler surface, and suppose that (M, J) is the minimal resolution of a surface singularity.

Corollary, suggested by Cristiano Spotti:

Theorem B. Let (M^4, g, J) be an ALE scalarflat Kähler surface, and suppose that (M, J) is the minimal resolution of a surface singularity. Then $m(M, g) \leq 0$,

Corollary, suggested by Cristiano Spotti:

Theorem B. Let (M^4, g, J) be an ALE scalar-flat Kähler surface, and suppose that (M, J) is the minimal resolution of a surface singularity. Then $m(M, g) \leq 0$, with = iff g is Ricci-flat.

Corollary, suggested by Cristiano Spotti:

Theorem B. Let (M^4, g, J) be an ALE scalar-flat Kähler surface, and suppose that (M, J) is the minimal resolution of a surface singularity. Then $m(M, g) \leq 0$, with = iff g is Ricci-flat.

Note that minimality is essential here.

Corollary, suggested by Cristiano Spotti:

Theorem B. Let (M^4, g, J) be an ALE scalar-flat Kähler surface, and suppose that (M, J) is the minimal resolution of a surface singularity. Then $m(M, g) \leq 0$, with = iff g is Ricci-flat.

Note that minimality is essential here.

Non-minimal resolutions typically admit families of such metrics for which the mass can be continuously deformed from negative to positive.

Lemma.

Lemma. Let (M, g) be any ALE manifold

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$.

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map $H_c^2(M) \to H_{dR}^2(M)$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Here

$$H_c^p(M) := rac{\ker d : \mathcal{E}_c^p(M) o \mathcal{E}_c^{p+1}(M)}{d\mathcal{E}_c^{p-1}(M)}$$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Here

$$H_c^p(M) := \frac{\ker d : \mathcal{E}_c^p(M) \to \mathcal{E}_c^{p+1}(M)}{d\mathcal{E}_c^{p-1}(M)}$$

where

 $\mathcal{E}_c^p(M) := \{ \text{Smooth, compactly supported } p \text{-forms on } M \}.$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map $H_c^2(M) \to H_{dR}^2(M)$

 $rac(W) \rightarrow radR$

is an isomorphism.

If just one end:

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

If just one end:

Compactify M as $\overline{M} = M \cup (S^{n-1}/\Gamma)$.

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

If just one end:

Compactify M as $\overline{M} = M \cup (S^{n-1}/\Gamma)$.

Then $H_c^p(M) = H^p(\overline{M}, \partial \overline{M}).$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

If just one end:

Compactify M as $\overline{M} = M \cup (S^{n-1}/\Gamma)$.

Then
$$H_c^p(M) = H^p(\overline{M}, \partial \overline{M}).$$

Exact sequence of pair:

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

If just one end:

Compactify M as $\overline{M} = M \cup (S^{n-1}/\Gamma)$.

Then
$$H_c^p(M) = H^p(\overline{M}, \partial \overline{M}).$$

Exact sequence of pair:

$$H^1(\partial \overline{M}) \to H^2(\overline{M}, \partial \overline{M}) \to H^2(\overline{M}) \to H^2(\partial \overline{M})$$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

If just one end:

Compactify M as $\overline{M} = M \cup (S^{n-1}/\Gamma)$.

Then
$$H_c^p(M) = H^p(\overline{M}, \partial \overline{M}).$$

Exact sequence of pair:

$$H^{1}(S^{n-1}/\Gamma) \to H^{2}(M) \to H^{2}(M) \to H^{2}(S^{n-1}/\Gamma)$$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Definition.

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Definition. If (M, g, J) is any ALE Kähler manifold,

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Definition. If (M, g, J) is any ALE Kähler manifold, we will use

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Definition. If (M, g, J) is any ALE Kähler manifold, we will use

$$\clubsuit: H^2_{dR}(M) \to H^2_c(M)$$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Definition. If (M, g, J) is any ALE Kähler manifold, we will use

$$\clubsuit: H^2_{dR}(M) \to H^2_c(M)$$

to denote the inverse of the natural map

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Definition. If (M, g, J) is any ALE Kähler manifold, we will use

$$\clubsuit: H^2_{dR}(M) \to H^2_c(M)$$

to denote the inverse of the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

Lemma. Let (M, g) be any ALE manifold of real dimension $n \geq 4$. Then the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

is an isomorphism.

Definition. If (M, g, J) is any ALE Kähler manifold, we will use

$$\clubsuit: H^2_{dR}(M) \to H^2_c(M)$$

to denote the inverse of the natural map

$$H_c^2(M) \to H_{dR}^2(M)$$

induced by the inclusion of compactly supported smooth forms into all forms.

Theorem C.

Theorem C. Any ALE Kähler manifold (M, g, J)

Theorem C. Any ALE Kähler manifold (M, g, J) of complex dimension m

$$m(M,g) = +$$

$$m(M,g) = + \int_{M} s_g d\mu_g$$

$$m(M,g) = + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M s_g d\mu_g$$

$$\mathbf{m}(M,g) = -\frac{\langle \mathbf{A}(c_1), [\omega]^{m-1} \rangle}{4(2m-1)\pi^m} \int_M \mathbf{s}_g d\mu_g$$

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\boldsymbol{\omega}]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_{M} \mathbf{s}_g d\mu_g$$

$$\begin{split} \textit{m}(\textit{M},g) &= -\frac{\langle \clubsuit(\textit{c}_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_{\textit{M}} \textit{s}_g d\mu_g \\ \textit{where} \end{split}$$

Theorem C. Any ALE Kähler manifold (M, g, J) of complex dimension m has mass given by

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M \mathbf{s}_g d\mu_g$$
 where

 \bullet s = scalar curvature;

$$m(M,g) = -\frac{\langle \mathbf{A}(c_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M s_g d\mu_g$$
 where

- \bullet s = scalar curvature;
- $d\mu = metric\ volume\ form;$

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M \mathbf{s}_g d\mu_g$$
 where

- \bullet s = scalar curvature;
- $d\mu = metric\ volume\ form;$
- $c_1 = c_1(M, J) \in H^2(M)$ is first Chern class;

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M \mathbf{s}_g d\mu_g$$
 where

- \bullet s = scalar curvature;
- $d\mu = metric\ volume\ form;$
- $c_1 = c_1(M, J) \in H^2(M)$ is first Chern class;
- $[\omega] \in H^2(M)$ is Kähler class of (g, J); and

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M \mathbf{s}_g d\mu_g$$
 where

- \bullet s = scalar curvature;
- $d\mu = metric\ volume\ form;$
- $c_1 = c_1(M, J) \in H^2(M)$ is first Chern class;
- $[\omega] \in H^2(M)$ is Kähler class of (g, J); and
- ullet $\langle \ , \ \rangle$ is pairing between $H_c^2(M)$ and $H^{2m-2}(M)$.

$$m(M,g) = -\frac{\langle \mathbf{A}(c_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M s_g d\mu_g$$

$$\frac{4\pi^m(2m-1)}{(m-1)!} \mathbf{m}(M,g) = -\frac{4\pi}{(m-1)!} \langle \mathbf{A}(\mathbf{c}_1), [\omega]^{m-1} \rangle + \int_M \mathbf{s}_g d\mu_g$$

For a compact Kähler manifold (M^{2m}, g, J) ,

$$\int_{M} s_{g} d\mu_{g} = \frac{4\pi}{(m-1)!} \langle c_{1}, [\omega]^{m-1} \rangle$$

For a compact Kähler manifold (M^{2m}, g, J) ,

$$0 = -\frac{4\pi}{(m-1)!} \langle c_1, [\omega]^{m-1} \rangle + \int_M s_g d\mu_g$$

For an ALE Kähler manifold (M^{2m}, g, J) ,

$$m(M,g) = -\frac{\langle \mathbf{A}(c_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M s_g d\mu_g$$

For an ALE Kähler manifold (M^{2m}, g, J) ,

$$\frac{4\pi^m(2m-1)}{(m-1)!} \mathbf{m}(M,g) = -\frac{4\pi}{(m-1)!} \langle \mathbf{A}(c_1), [\omega]^{m-1} \rangle + \int_M s_g d\mu_g$$

So the mass is a "boundary correction" to the topological formula for the total scalar curvature.

$$\mathbf{m}(\mathbf{M},g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\boldsymbol{\omega}]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_{\mathbf{M}} \mathbf{s}_g d\mu_g$$

Corollary. Any ALE scalar-flat Kähler manifold (M, g, J) of complex dimension m has mass given by

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\boldsymbol{\omega}]^{m-1} \rangle}{(2m-1)\pi^{m-1}}.$$

Corollary. Any ALE scalar-flat Kähler manifold (M, g, J) of complex dimension m has mass given by

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\boldsymbol{\omega}]^{m-1} \rangle}{(2m-1)\pi^{m-1}}.$$

So **Theorem** A is an immediate consequence!

Theorem D (Positive Mass Theorem).

Theorem D (Positive Mass Theorem). Any AE $K\ddot{a}hler \ manifold \ with$

Theorem D (Positive Mass Theorem). Any AE Kähler manifold with non-negative scalar curvature

 $AE \& K\"{a}hler \& s \ge 0 \implies m(M,g) \ge 0.$

$$AE \& K\"{a}hler \& s \ge 0 \implies m(M, g) \ge 0.$$
 $Moreover, m = 0 \iff$

 $AE \& K\"{a}hler \& s \ge 0 \implies m(M,g) \ge 0.$

Moreover, $m = 0 \iff (M, g)$ is Euclidean space.

$$AE \& K\"{a}hler \& s \ge 0 \implies m(M,g) \ge 0.$$
 Moreover, $m = 0 \iff (M,g)$ is Euclidean space.

Proof actually shows something stronger!

Theorem E (Penrose Inequality).

Theorem E (Penrose Inequality). Let (M^{2m}, g, J)

Theorem E (Penrose Inequality). Let (M^{2m}, g, J) be an AE Kähler manifold

Theorem E (Penrose Inequality). Let (M^{2m}, g, J) be an AE Kähler manifold with scalar curvature $s \geq 0$.

canonical \iff holomorphic section of $K = \Lambda^{m,0}$

canonical \Longrightarrow Poincaré dual to $-c_1$.

Theorem E (Penrose Inequality). Let (M^{2m}, g, J) be an AE Kähler manifold with scalar curvature $s \geq 0$. Then (M, J) carries a canonical divisor D that is expressed as a sum $\sum_{j} \mathbf{n}_{j} D_{j}$

Theorem E (Penrose Inequality). Let (M^{2m}, g, J) be an AE Kähler manifold with scalar curvature $s \geq 0$. Then (M, J) carries a canonical divisor D that is expressed as a sum $\sum_{j} \mathbf{n}_{j} D_{j}$ of compact complex hypersurfaces with positive integer coefficients,

$$m(M,g) \ge Vol(D_j)$$

$$m(M,g) \ge \sum_{j=1}^{n} Vol(D_j)$$

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{j=1}^{n} \operatorname{Vol}(D_j)$$

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{\mathbf{n}_j} Vol(D_j)$$

 $with = \iff$

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{j=1}^{n} \operatorname{Vol}(D_j)$$

 $with = \iff (M, g, J) \text{ is scalar-flat K\"{a}hler.}$

$$\mathbf{m}(\mathbf{M},g) = \frac{\langle \mathbf{A}(-\mathbf{c}_1), [\boldsymbol{\omega}]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_{\mathbf{M}} \mathbf{s}_g d\mu_g$$

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{j=1}^{n} \operatorname{Vol}(D_j)$$

 $with = \iff (M, g, J) \text{ is scalar-flat K\"{a}hler.}$

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{j=1}^{n} \operatorname{Vol}(D_j)$$

 $with = \iff (M, g, J) \text{ is scalar-flat K\"{a}hler.}$

In complex dimension ≥ 3 , only requires Chruściel metric fall-off.

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{\mathbf{n}_j} Vol(D_j)$$

 $with = \iff (M, g, J) \text{ is scalar-flat K\"{a}hler.}$

In complex dimension ≥ 3 , only requires Chruściel metric fall-off.

$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad s \in L^1$$

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{j=1}^{n} \operatorname{Vol}(D_j)$$

 $with = \iff (M, g, J) \text{ is scalar-flat K\"{a}hler.}$

In complex dimension ≥ 3 , only requires Chruściel metric fall-off.

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{\mathbf{n}_j} Vol(D_j)$$

 $with = \iff (M, g, J)$ is scalar-flat Kähler.

In complex dimension ≥ 3 , only requires Chruściel metric fall-off.

In complex dimension 2, we needed stronger fall-off

$$g_{jk} - \delta_{jk} \in C^{2,\alpha}_{-\tau}, \quad \tau > \frac{n-2}{2}.$$

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{j=1}^{n} \operatorname{Vol}(D_j)$$

 $with = \iff (M, g, J) \text{ is scalar-flat K\"{a}hler.}$

But new proof in today's ArXiv posting only requires Chruściel fall-off in any dimension.

$$m(M,g) \ge \frac{(m-1)!}{(2m-1)\pi^{m-1}} \sum_{\mathbf{n}_j} Vol(D_j)$$

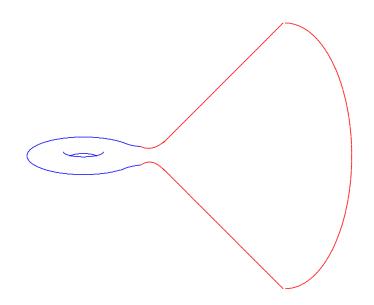
 $with = \iff (M, g, J) \text{ is scalar-flat K\"{a}hler.}$

But new proof in today's ArXiv posting only requires Chruściel fall-off in any dimension.

$$g_{jk} = \delta_{jk} + O(|x|^{1 - \frac{n}{2} - \varepsilon})$$
$$g_{jk,\ell} = O(|x|^{-\frac{n}{2} - \varepsilon}), \quad s \in L^1$$

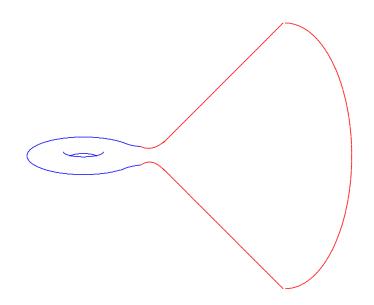
$$m(M,g) = -\frac{\langle \mathbf{A}(c_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M s_g d\mu_g$$

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\boldsymbol{\omega}]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_{M} \mathbf{s}_g d\mu_g$$



$$m(M,g) = -\frac{\langle \mathbf{A}(c_1), [\omega]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_M s_g d\mu_g$$

$$m(M,g) = -\frac{\langle \mathbf{A}(\mathbf{c}_1), [\boldsymbol{\omega}]^{m-1} \rangle}{(2m-1)\pi^{m-1}} + \frac{(m-1)!}{4(2m-1)\pi^m} \int_{M} \mathbf{s}_g d\mu_g$$



End, Part I